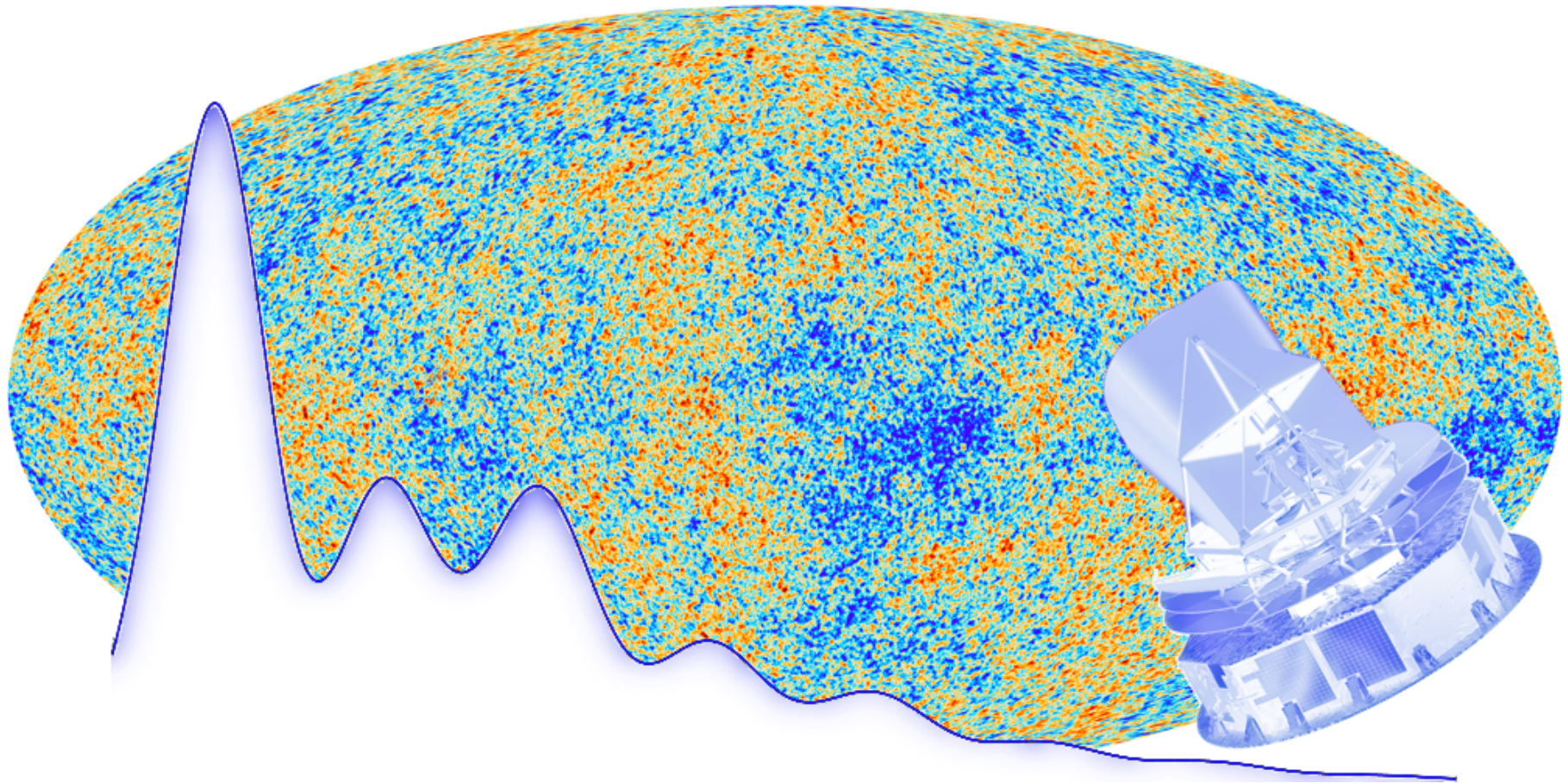


# The Cosmological Legacy of Planck

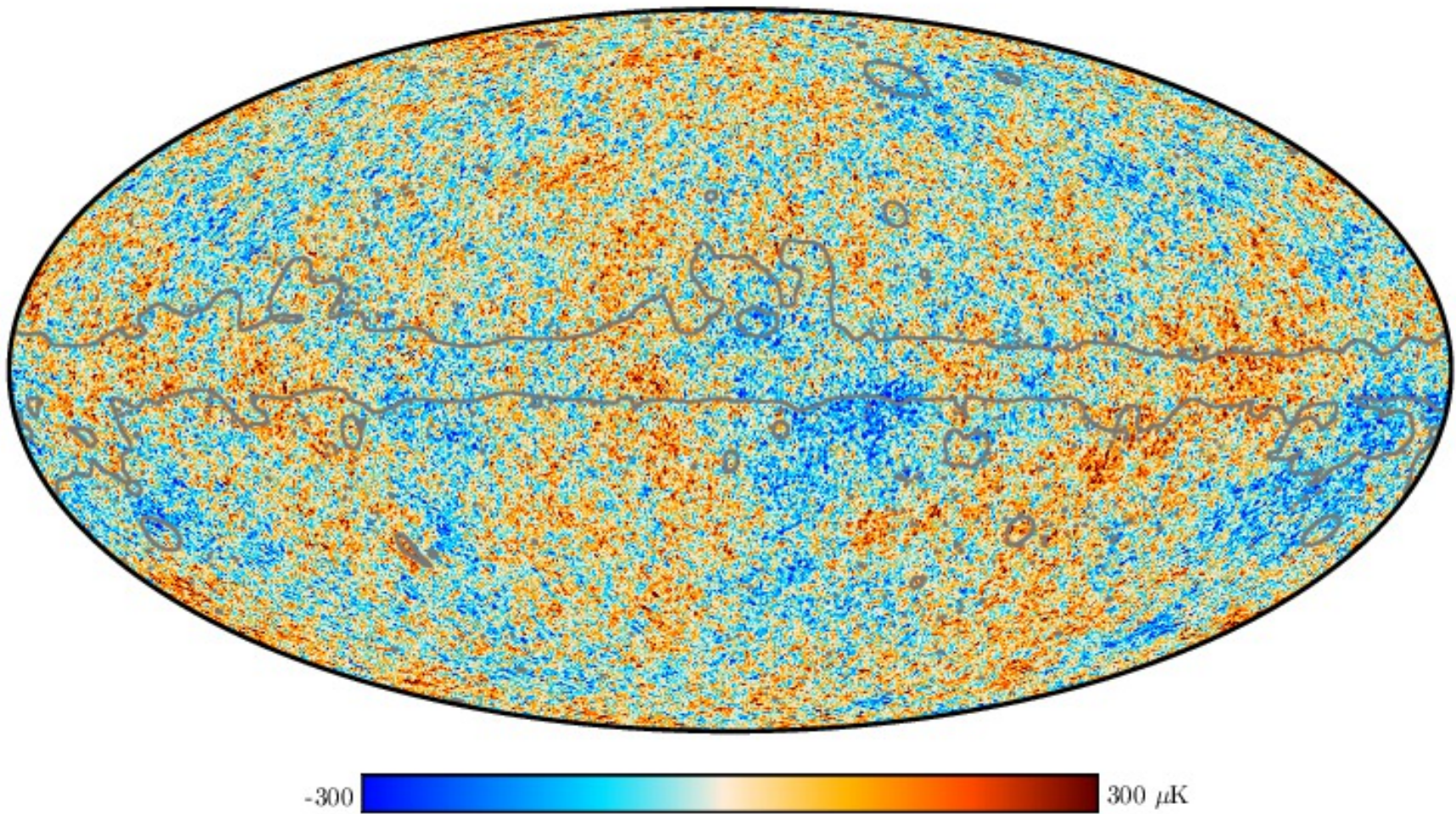


Martin White  
Berkeley.

Figs. courtesy V. Pettorino

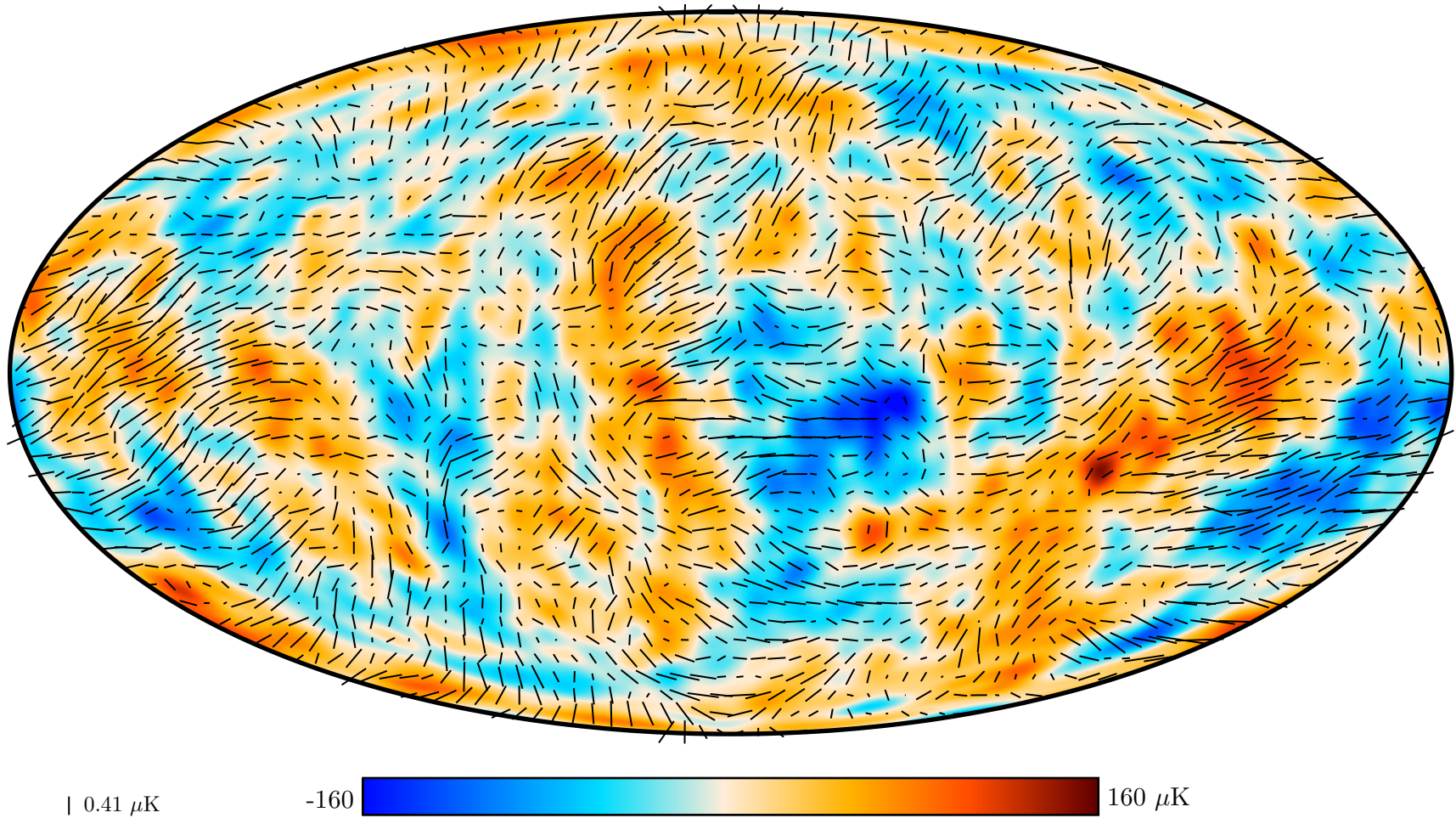


# CMB map

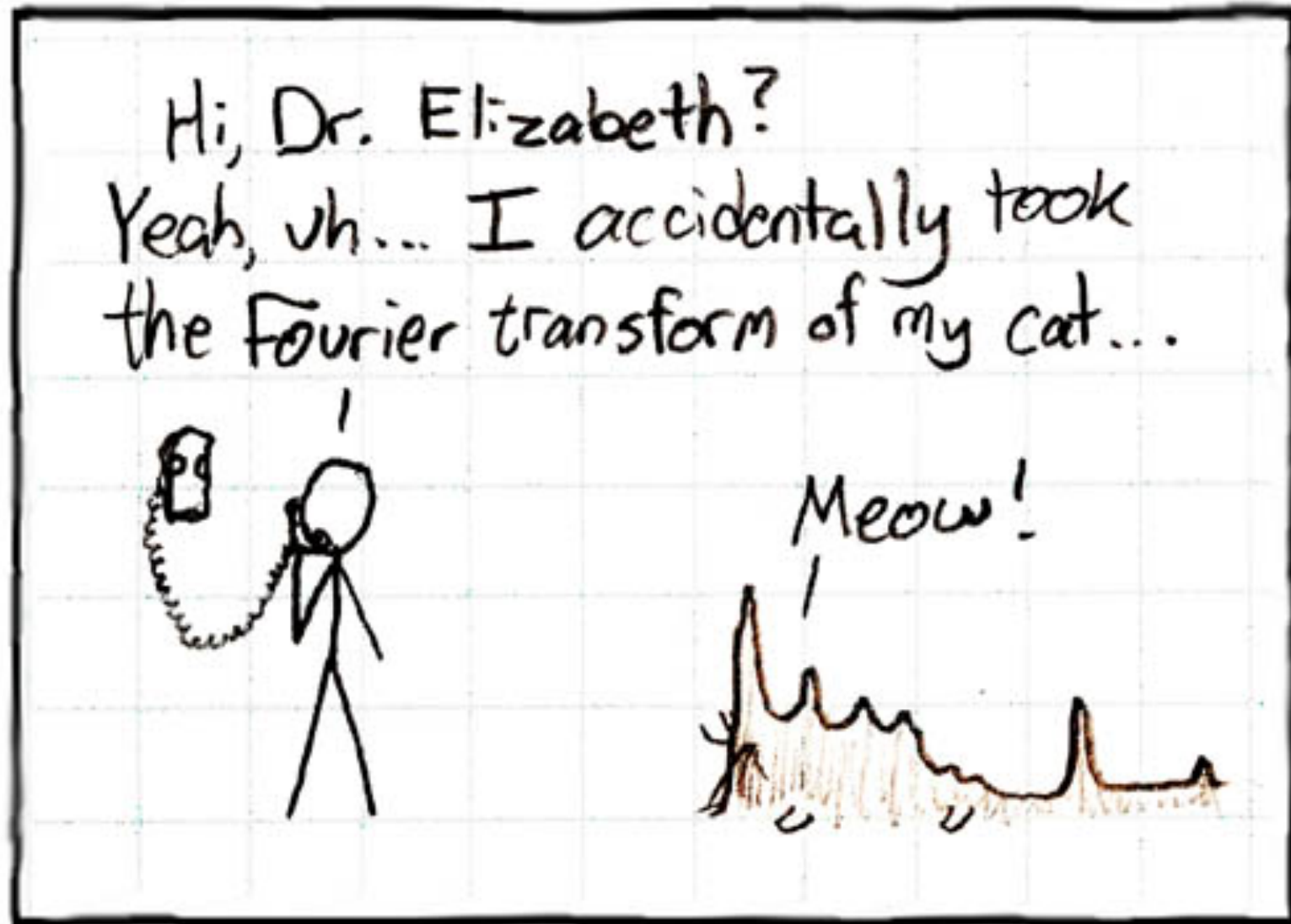




# CMB map: smoothed + polarization



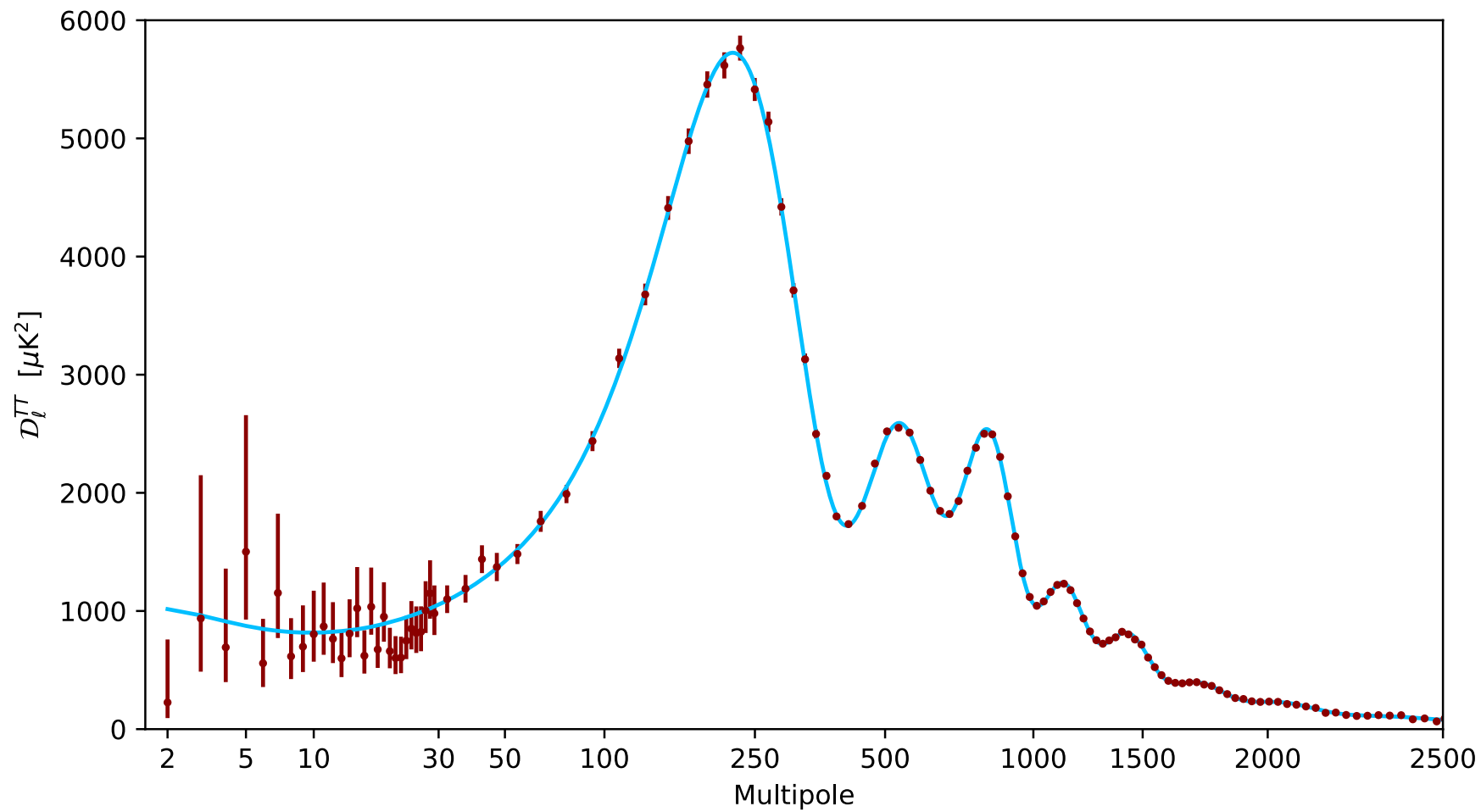
## Power spectrum ...



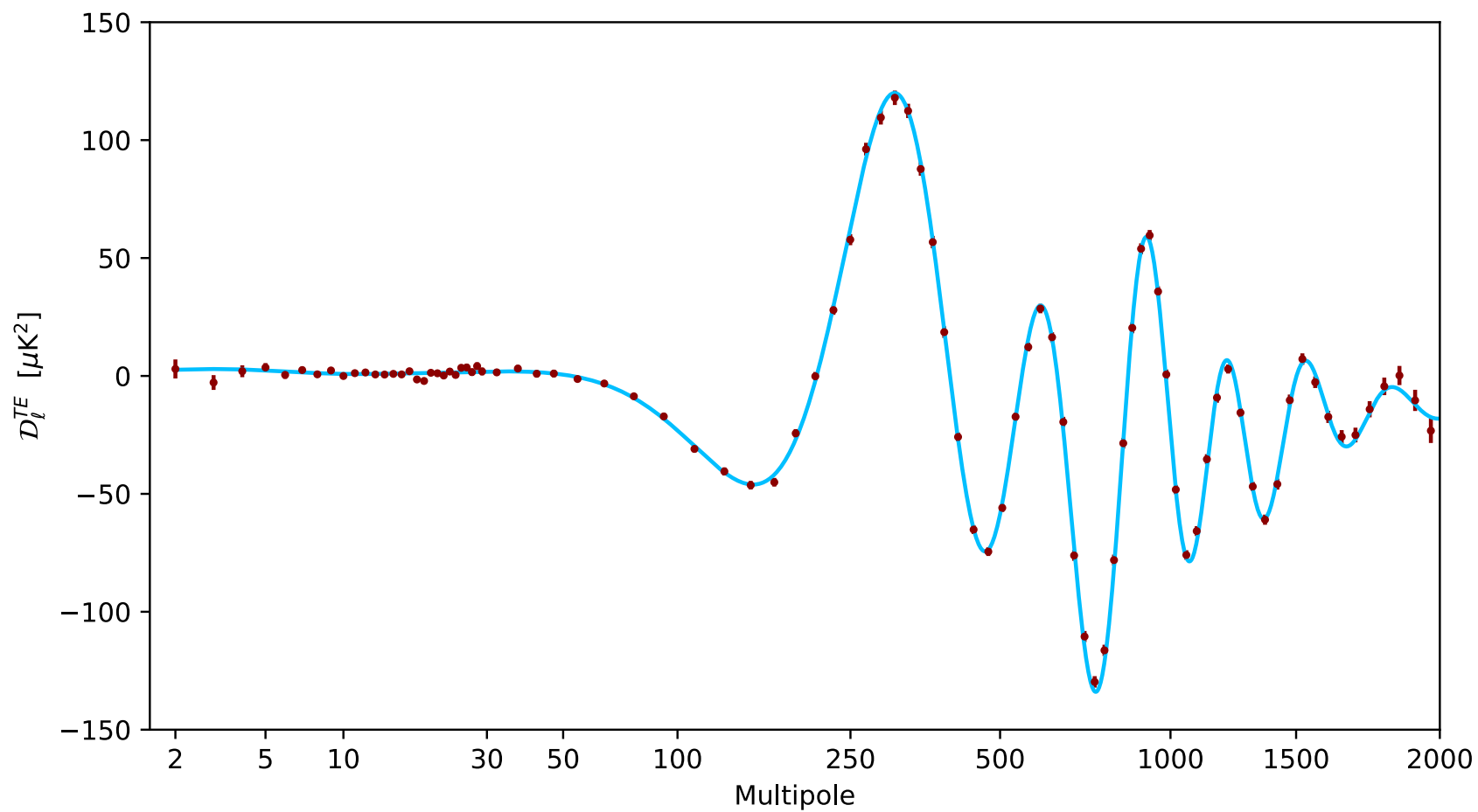
[xkcd.com/26](http://xkcd.com/26)



# The angular power spectrum

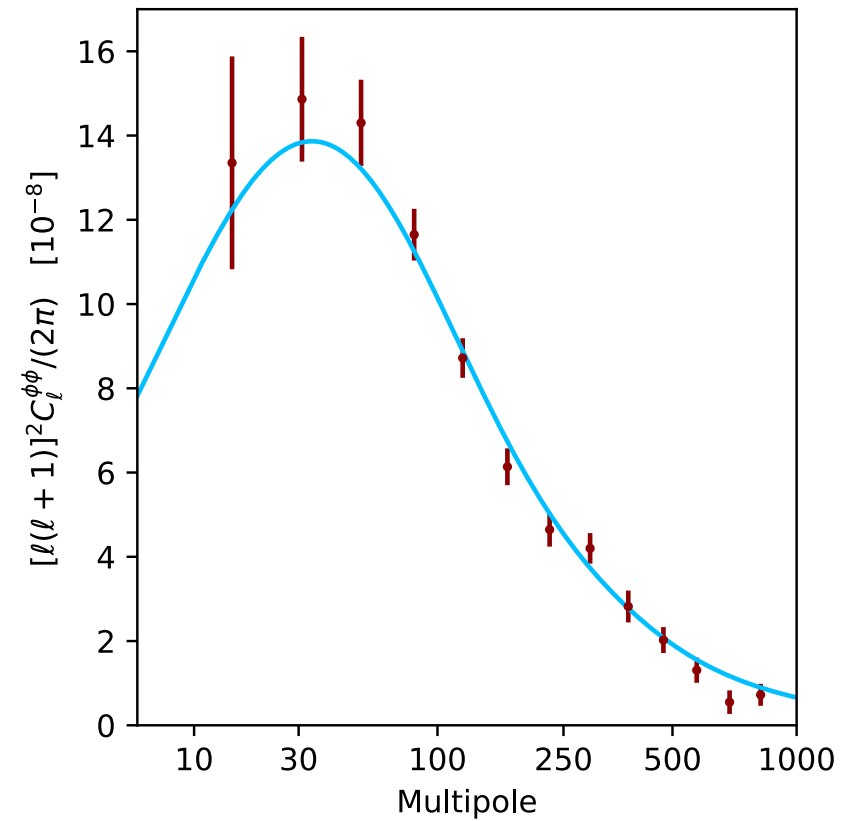
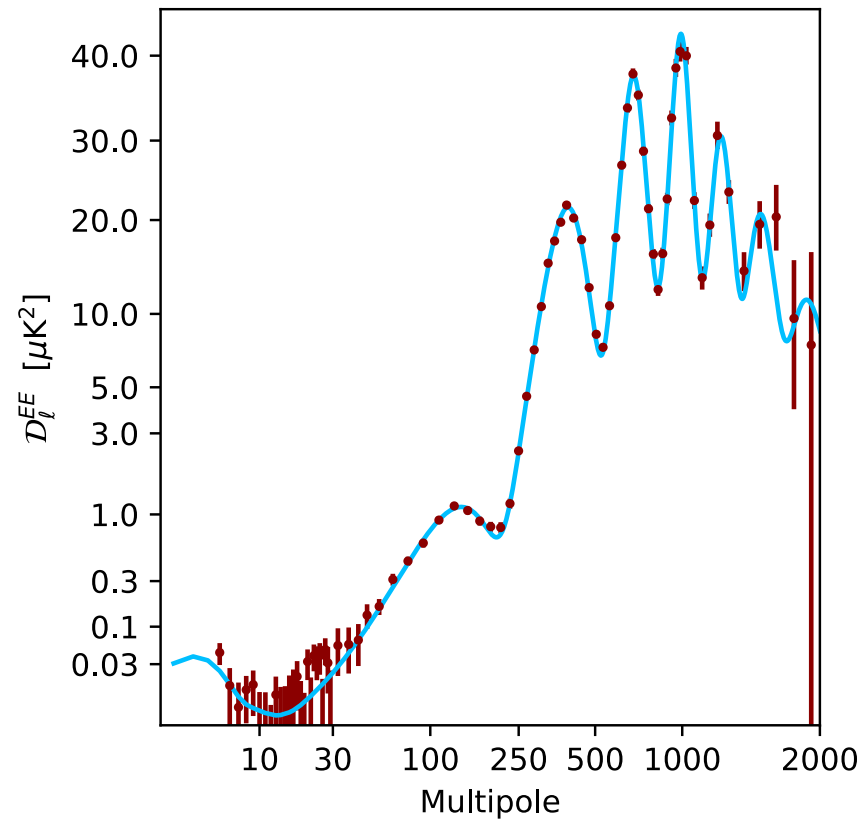


# Polarization-Temperature





# Polarization and lensing



# Data compression!

- We find that a simple, 6 parameter model fits the data extremely well.
  - Data compression: trillions of bits of data are compressed to billions of measurements at 9 frequencies, then tens of millions of modes are compressed to thousands of multipoles which are compressed to 6 cosmological parameters!
  - With no evidence for a 7<sup>th</sup>.
- For the “base model” the CMB determines all of the parameters, on its own, with exceptional accuracy.
  - If we include polarization, best determined parameter is 0.03%.
  - Only 1 parameter not determined to better than 1%.



# Planck(-only) base $\Lambda$ CDM model

Parameter	Description	Value
$\omega_b$	Baryon density	$0.02237 \pm 0.00015$
$\omega_c$	Cold dark matter density	$0.1200 \pm 0.0012$
$100\theta_{MC}$	Angular size of acoustic scale	$1.04092 \pm 0.00031$
$\tau$	Optical depth to Thomson scattering	$0.0544 \pm 0.0073$
$\ln(10^{10}A_s)$	Observed fluctuation amplitude	$3.044 \pm 0.014$
$n_s$	Slope of primordial power spectrum (spectral index)	$0.9649 \pm 0.0042$
<hr/>		
$H_0$ (km/s/Mpc)	Expansion rate of Universe	$67.36 \pm 0.54$
$\sigma_8$	Amplitude of fluctuations in matter today	$0.8111 \pm 0.006$

And my favorite derived parameter:  $k_{eq} = 0.01038 \pm 0.00008 \text{ Mpc}^{-1}$

# **Inflationary $\Lambda$ CDM: a great phenomenological model**

- Model has withstood incredible increases in data quality over the last 3+ decades.
  - Model predictions for anisotropy spectra were quite specific.
  - Many extensions/variants now highly constrained.
- Puzzling contents:
  - Neutrino masses are  $O(100\text{meV})$  not  $O(100\text{GeV})$ .
  - $\omega_B^{\text{BBN}} = \omega_B^{\text{CMB}} \sim \omega_M$  ; DE smooth & “turns on” rapidly today.
  - No sign of “extra” relativistic species or spatial curvature ...
- Model connects high-energy physics to cosmology and “explains” 14Gyr of cosmic evolution, but our understanding is “highly incomplete”!



# Planck & Inflation

- CMB quickly established early Universe origin of perturbations.
- Planck has had a huge impact on inflationary model building!
- A large number of “popular” models now ruled out.
- The simplest models of inflation predict ...

A spatially flat Universe	$\Omega_K = 0.0007 \pm 0.0019$
with <i>nearly</i> scale-invariant (red) spectrum of density perturbations	$0.967 \pm 0.004$
which is almost a power-law	$dn_s/d\ln k = -0.0042 \pm 0.0067$
dominated by scalar perturbations	$r_{0.002} < 0.1$ (95%; $< 0.07 + \text{BKP}$ )
which are Gaussian	$f_{\text{NL}} = -0.9 \pm 5.1 \sim 0$
and adiabatic	$\alpha_{-1} = 0.00013 \pm 0.00037$
with negligible topological defects	$f_{\text{NG}} < 0.01$ (95%)

# Inflationary models

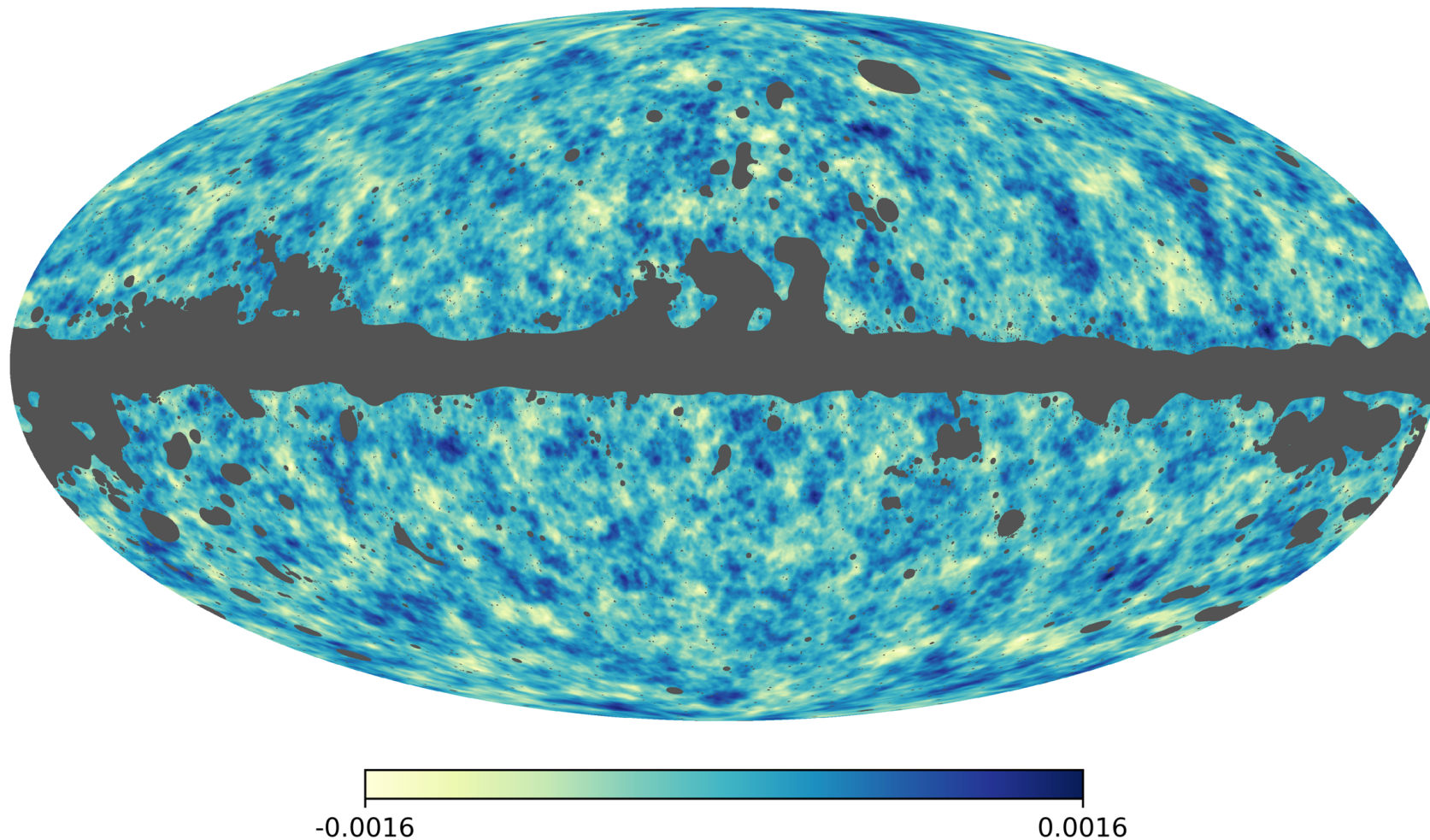
- Coherence of peaks, sign of TE
  - Early Universe origin of perturbations
- $\Omega_K \sim 0$ : duration of slow-roll not fine tuned.
- Primordial  $P(k)$  well approximated by power-law.
  - Inflaton rolls down a featureless, nearly flat potential.
- No isocurvature modes: 1 d.o.f.
- Scalar modes dominate by 1 order of magnitude.
  - Models with  $r \sim (1-n_s)$  severely limited.
  - Models with  $r \sim (1-n_s)^2$  require next-gen technology to limit.
  - Models with  $r \ll (1-n_s)^2$  out of reach of foreseeable technology.
- Surviving models have  $V' \sim 0$  and  $V'' < 0$ 
  - special point in potential.



# CMB lensing

- Photons from the CMB are deflected on their way to us by the potentials due to large-scale structure.
- Gives sensitivity to the “low  $z$ ” Universe.
  - Allows us to break some degeneracies from purely within the CMB dataset.
  - Provides a cross-check on the paradigm: are the structures we infer at  $z \sim 2$  consistent with the “initial conditions” measured at  $z \sim 1,000$ ? [After  $10^3$  growth:  $A_{\phi\phi} = 0.997 \pm 0.03$ ]
- Provides a map, over the whole sky, of the (projected) mass back to the surface of last-scattering (98% of the way to the horizon).

# Lensing deflection (E-mode)

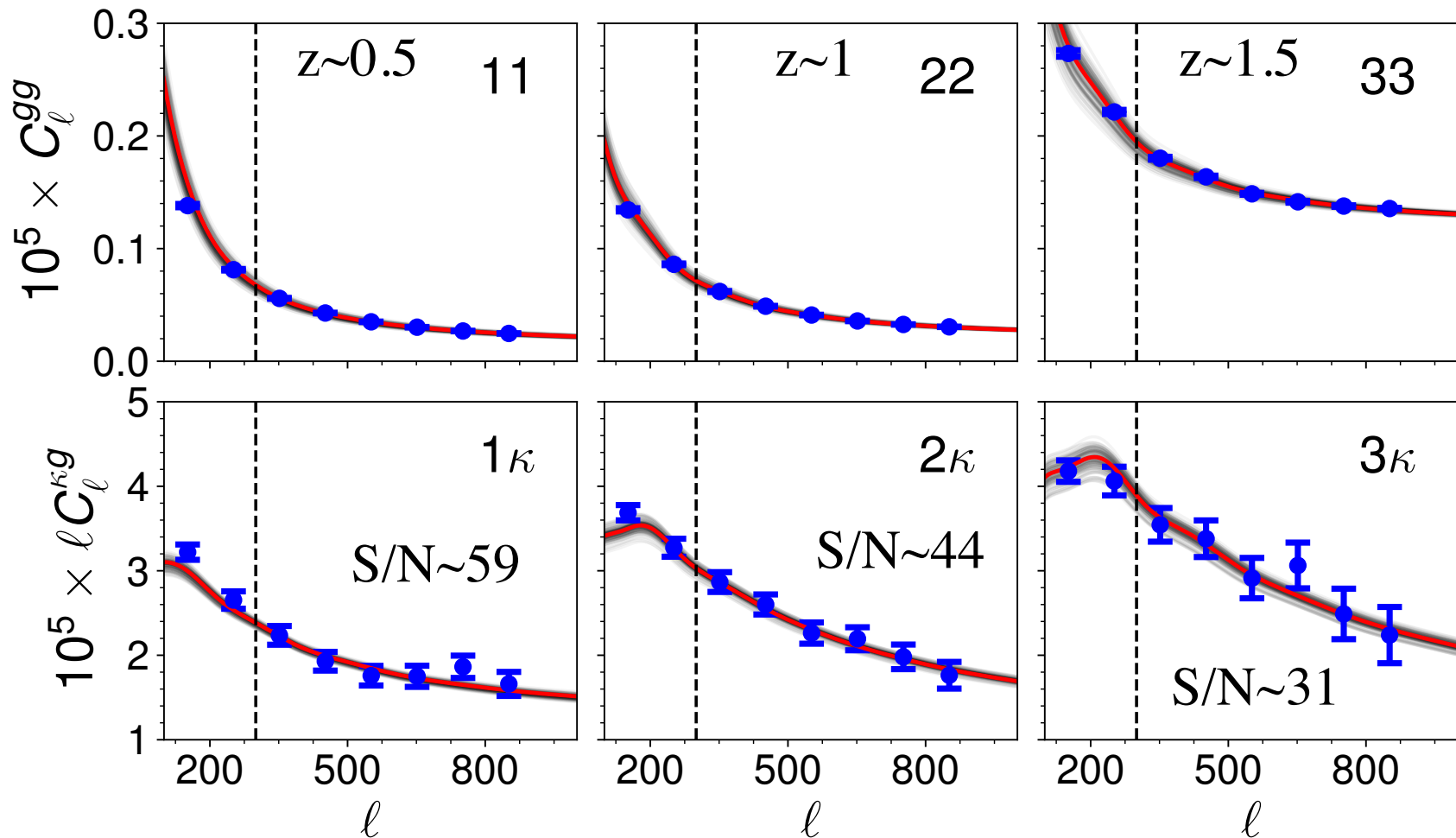


Lensing now measured at  $>40\sigma$

## A lensing first ...

- ✧ Planck was definitely not the first experiment to
  - ✧ Measure gravitational lensing
  - ✧ ... by large scale structure
  - ✧ ... of the CMB
- ✧ However it was the first to do it over a significant fraction of the sky with enough S/N to drive fits and provide a sharp test of the theory.
- ✧ A “coming of age” story for CMB lensing.
- ✧ Planck ushers in a new era of CMB studies of the “intermediate” redshift Universe, synergistically with DES, WISE, HSC, DESI, PFS, LSST, Euclid, WFIRST, ...

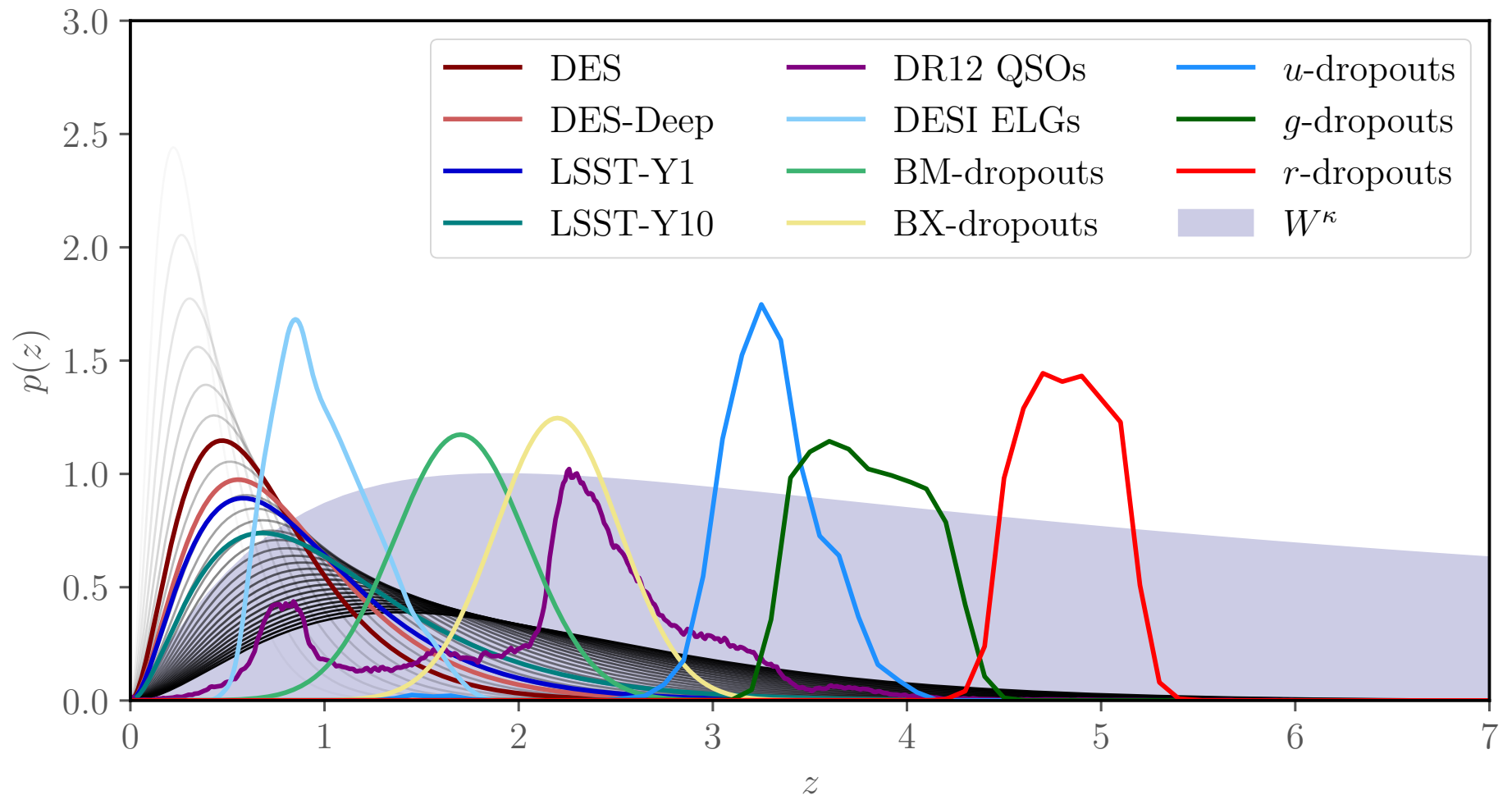
# Cross-correlation with (un)WISE galaxies



Krolewski, Ferraro++



# Tomographic decomposition of lensing kernel



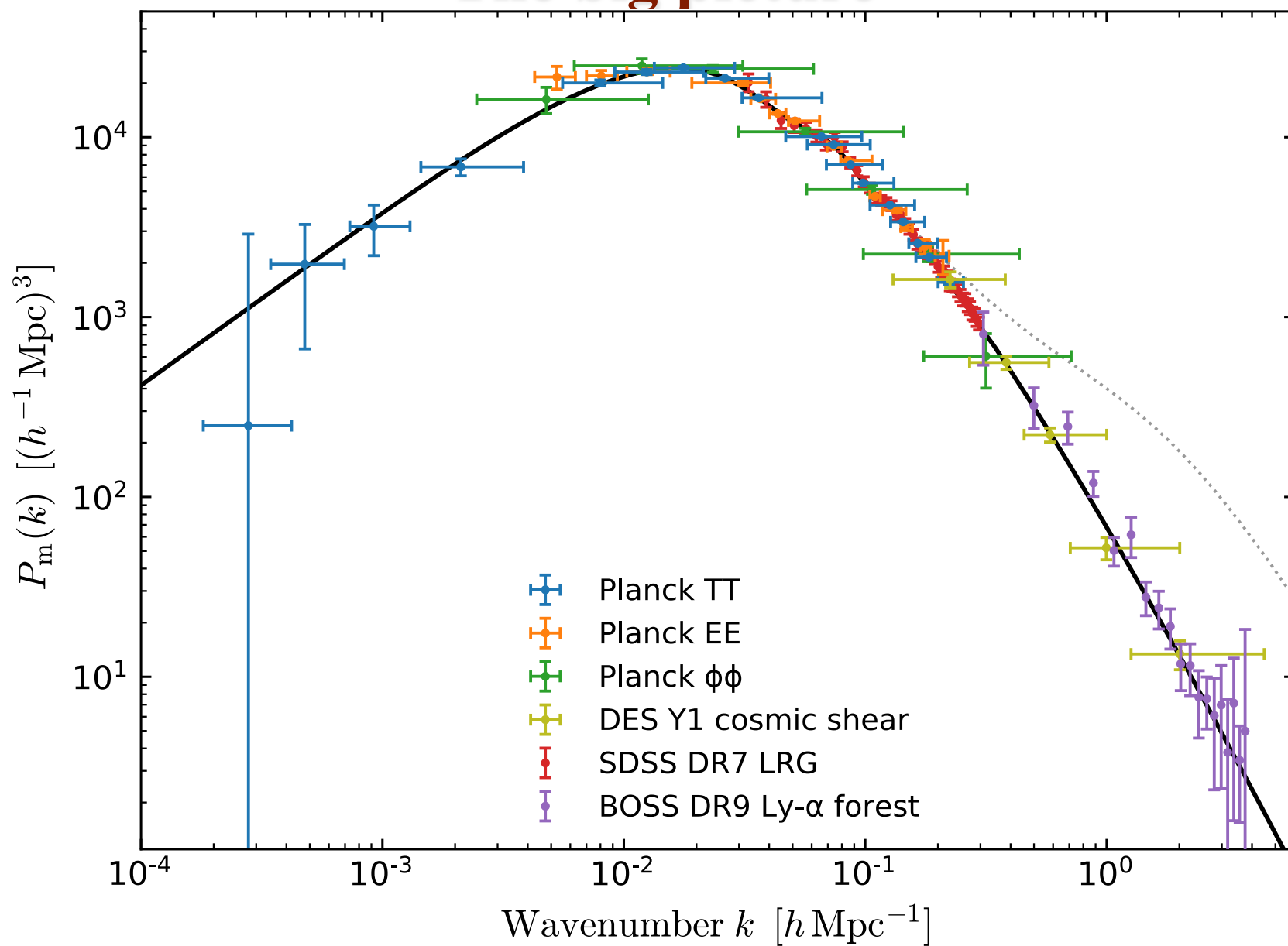
Wilson & White (2019)

## CMB + LSS

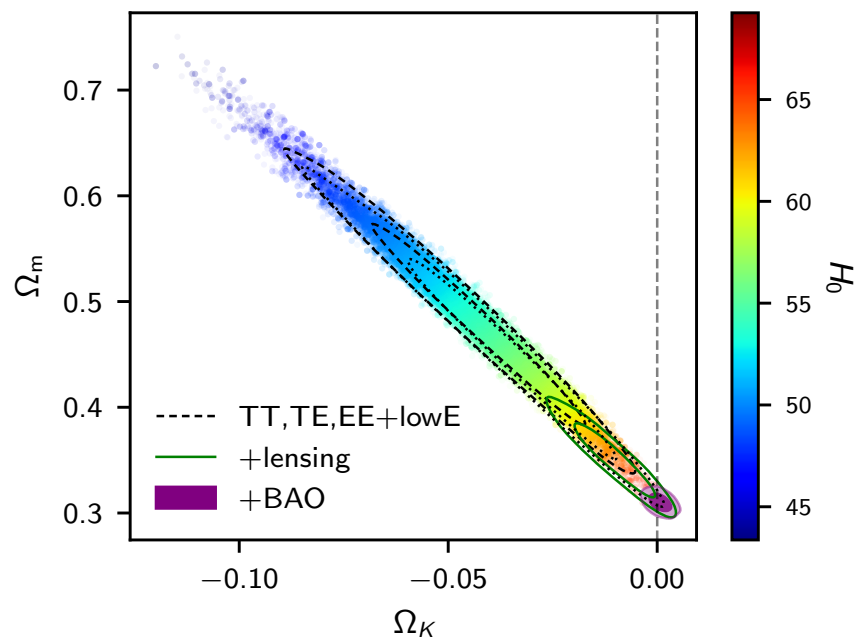
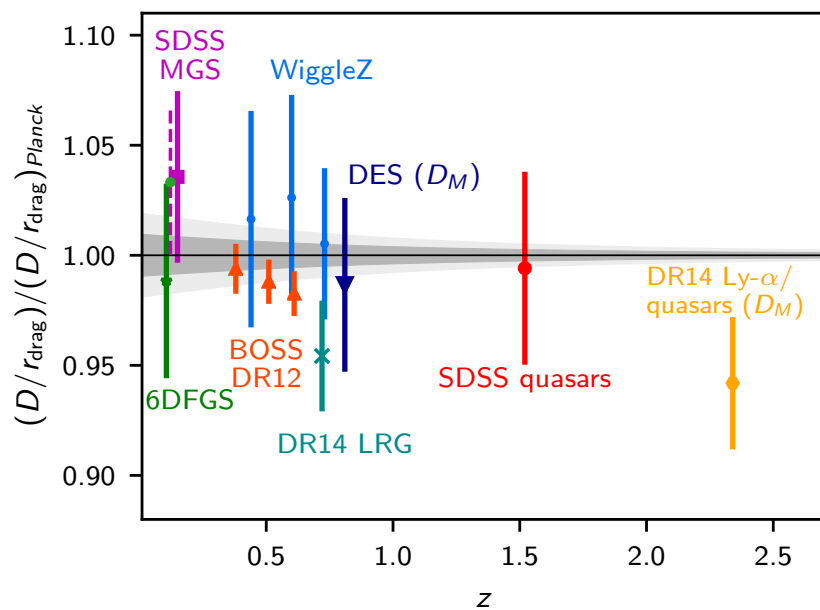
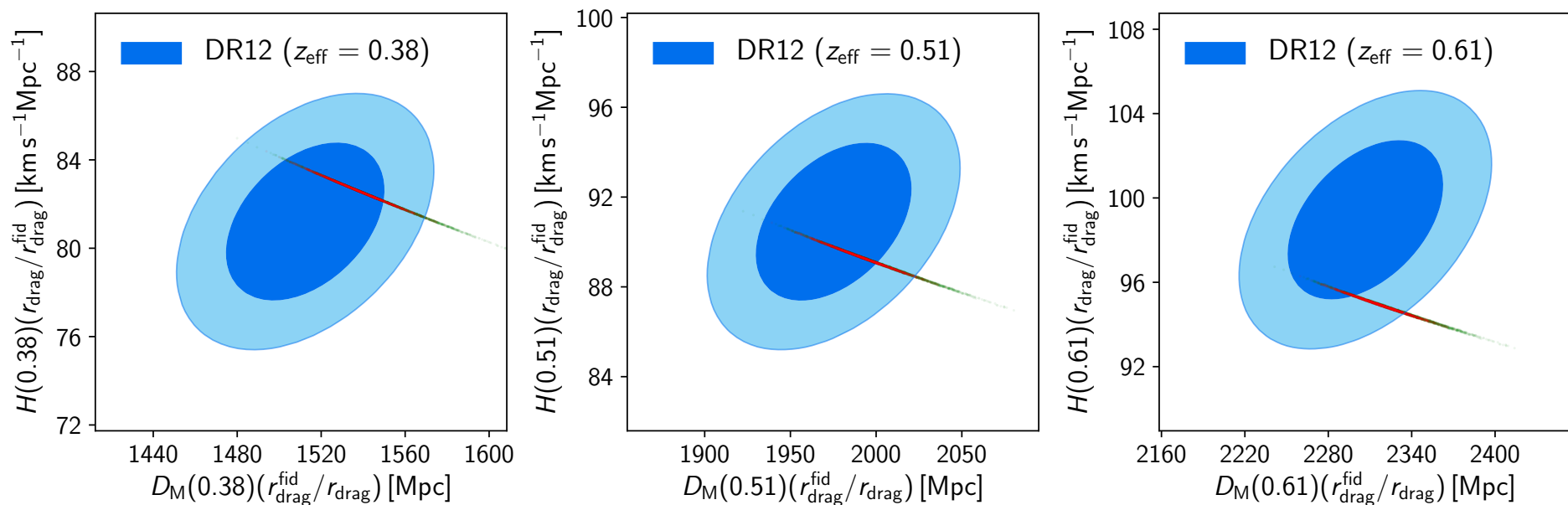
- By cementing the gravitational instability paradigm and measuring the ICs and parameters, Planck sets the framework for LSS.
- Planck precisely determines many of the key parameters for large-scale structure (in  $\Lambda$ CDM):
  - $k_{\text{eq}} = 0.01038 \pm 0.00008 \text{ Mpc}^{-1}$ ,  $\sigma_8(z=2) = 0.3211 \pm 0.0009$
- Planck calibrates the “standard fluctuation spectrum”.
  - Sets the scale and level of inhomogeneity in the Universe.
  - Governs structure formation, galaxy formation, etc.

Early on, the fields of LSS and CMB were tightly coupled. With time they grew apart and specialized. I think we are witnessing a re-coupling.

# The big picture



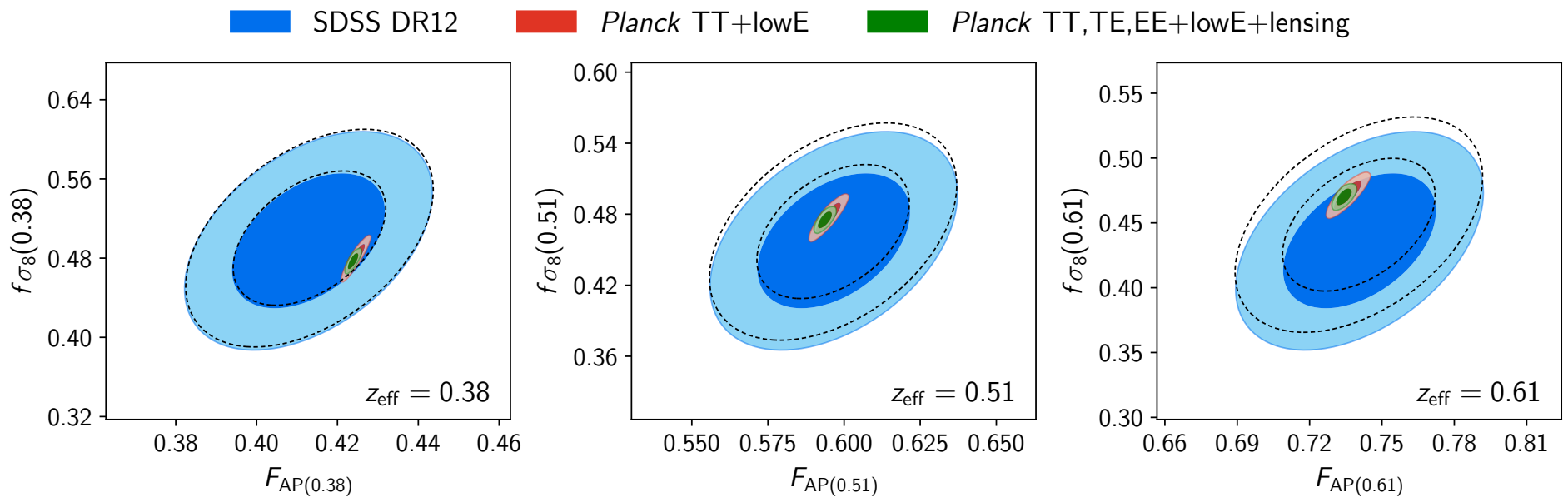
# Cosmic distance scale





# Redshift-space distortions: Planck vs BOSS

Growth rate ( $f\sigma_8$ ) vs. A-P parameter ( $F_{AP}$ )



(BOSS results have been marginalized over  $D_V$ :  
dashed lines show results conditioned on Planck)

## Just plain cool ...

- In 2013 Planck detected the motion of the Earth in the aberration of the measured CMB anisotropy.
  - Observed at  $>4\sigma$  in 2013 data.
- In 2015 we detected the impact of fluctuations in the 2K neutrino background!
  - Evidence for  $\nu$  background strong ( $N_{\text{eff}}=0$  ruled out @  $>10\sigma$ )
  - Now have exquisite detection of free-streaming of this component (measures of  $c_{\text{eff}}^2$  and  $c_{\text{vis}}^2$ ).
- In 2018 we measured the “gravitational slip” at  $z=1000$  to be  $1.004 \pm 0.007$ .
  - GR predicts it is 1.

# Conclusions

- Planck has “completed” the primary temperature anisotropy story begun by COBE.
  - Established acoustic physics as the “gold standard” probe.
- Impressive confirmation of the standard cosmological model.
  - Precise constraints on model and parameters.
  - Tight limits on deviations from base model.
  - Some indications of internal and external tensions, but with only modest\* statistical significance.
- Established CMB lensing as a competitive cosmological tool.
  - Synergies between large-scale structure and CMB are only growing in importance!
- The next decade will see a host of new facilities coming on line, enabling increasingly precise tests of our models ... building on the cosmological legacy of Planck.

*The End*