

Cosmology in 1 dimension

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“A man grows stale if he works all the time on insoluble problems, and a trip to the beautiful world of one dimension will refresh his imagination better than a dose of LSD.”

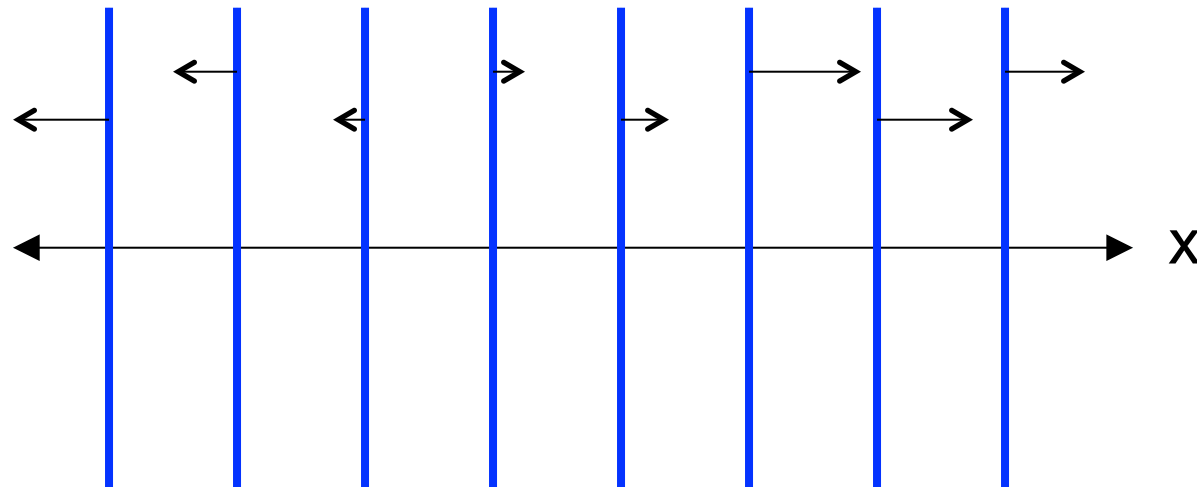
-- Freeman Dyson

Motivation

- There has been a great deal of work recently on (cosmological) perturbation theory.
 - New approaches, new resummation schemes, new renormalization techniques borrowed from QFT.
 - Growing appreciation of the uses and limitations of “standard perturbation theory” (SPT) and resummation schemes.
 - Understanding of RSD, BAO, SSC, beat-coupling, ...
- Want to understand these developments (and old ideas) better in a simple context:
- Collection of uniform, parallel, 2D sheets of matter.
 - Problem becomes 1 dimensional (plus time).
 - Significant (!) analytic simplification: can do SPT to ∞ order.
 - Easier to handle numerically with high dynamic range.
 - Many of the features of 3D have close 1D analogues.

The setup

(with $\Omega_m=1$ throughout)



$$P_{3D}(k_{\parallel}, \mathbf{k}_{\perp}) = (2\pi)^2 \delta^{(D)}(\mathbf{k}_{\perp}) P_{1D}(k_{\parallel})$$

Since the force on a particle due to a sheet is independent of the distance from the sheet, 1st order Lagrangian PT (Zeldovich) is exact until “sheet crossing” (can also show this analytically).

Power spectra and correlation functions

$$P_{\text{SPT}}^{1-\text{loop}}(k) = P_L(k) + \int_{-\infty}^{\infty} \frac{dk'}{2\pi} \left\{ 3 + 4 \frac{k-k'}{k'} + \left(\frac{k-k'}{k'} \right)^2 \right\} P_L(k') P_L(k-k') \\ - k^2 P_L(k) \int_{-\infty}^{\infty} \frac{dk'}{2\pi} \frac{P_L(k')}{k'^2}$$

$$\xi_{\text{SPT}}^{1-\text{loop}}(x) = \xi_L(x) + \overbrace{3\xi_L^2(x)}^{\text{growth}} + \overbrace{4\xi'_L(x) \int_x^{\infty} dx \xi_L(x)}^{\text{dilation}} + \overbrace{\frac{\sigma_{\text{eff}}^2}{2} \xi''_L(x)}^{\text{RMS displ.}} + \mathcal{O}(\xi_L^3)$$

Can look at the response of power spectra to long wavelength mode (through the gradient times the large-scale variance) and the modes themselves (super-sample covariance, beat coupling), shifts and broadening of the BAO peak, ...

PT issues

- Convergence rather slow.
 - Resummation schemes.
 - Beware symmetry breaking!
- Solutions only valid prior to “sheet” crossing.
- $P^{1\text{-loop}}(k)$ depends on high- k' , non-perturbative modes even at large scales.

Lagrangian theory

$$1 + \delta_{\text{LPT}}(x) = \int dq \delta^D[x - q - \Psi(q)]$$

$$\delta_{\text{LPT}}(k) = \int dq e^{-ikq} \left(e^{-ik\Psi(q)} - 1 \right)$$

But Ψ is just a Gaussian random variable ... know $\langle e^\Psi \rangle$

$$P_{\text{ZA}}(k) = \int dq e^{-ikq} \left(e^{-k^2 \sigma^2(q)/2} - 1 \right)$$

$$\sigma^2(q) = \langle [\Psi_{\text{ZA}}(0) - \Psi_{\text{ZA}}(q)]^2 \rangle = \int_0^\infty \frac{dk}{\pi} \frac{2 P_L(k)}{k^2} (1 - \cos[kq])$$

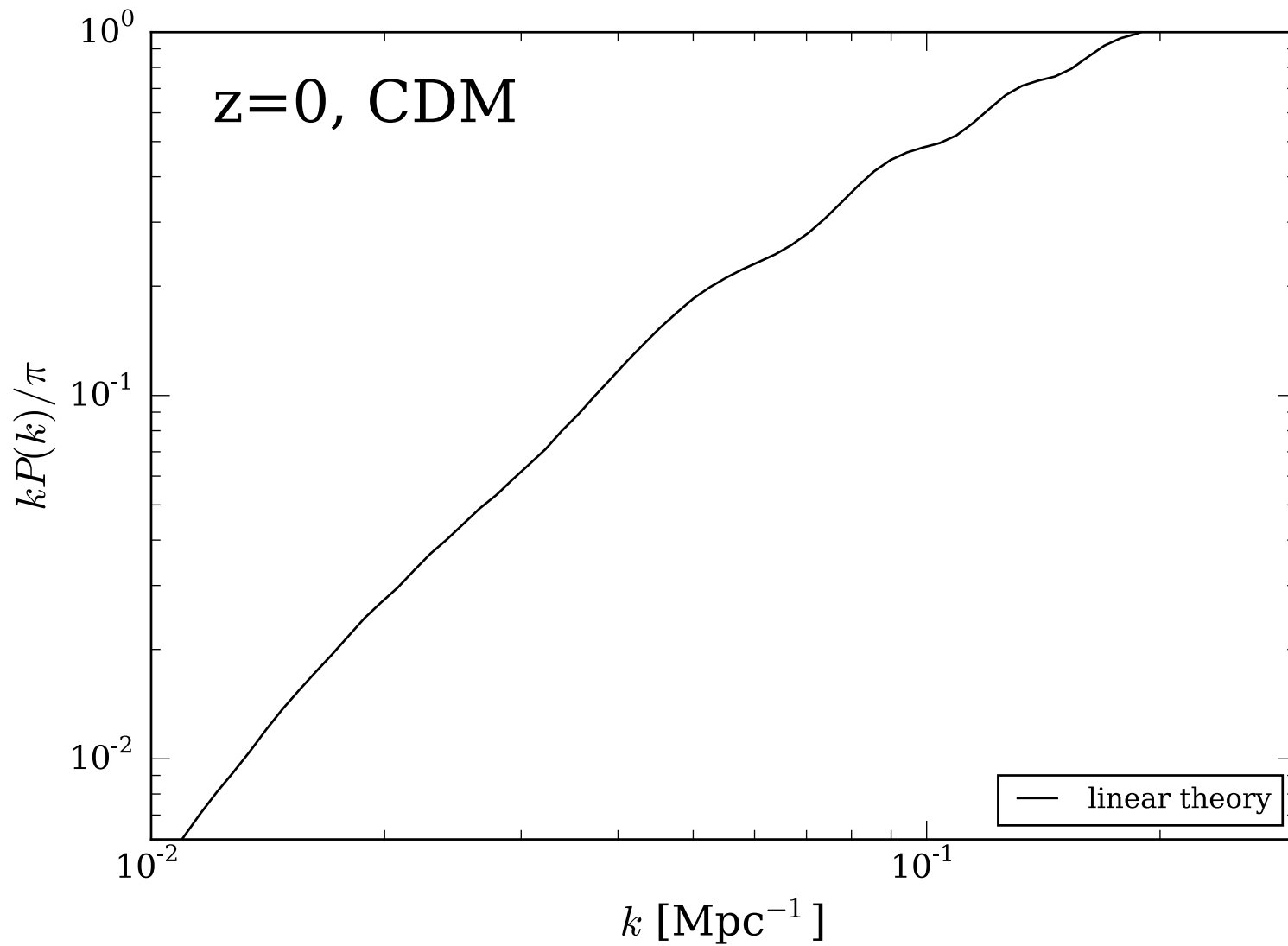
Can generate any order in PT!

$$\begin{aligned} P_{\text{LPT}}(k) &= \int dq e^{ikq} \left(-\frac{k^2}{2} \sigma^2(q) + \frac{k^4}{8} \sigma^4(q) + \dots \right), \\ &= P_L + \frac{1}{8} \int dq e^{ikq} \nabla_q^4 \sigma^4(q) + \mathcal{O}(P_L^3), \\ &= P_L + \frac{1}{8} \int dq e^{ikq} \left[6([\sigma^2]''')^2 + 8([\sigma^2]'[\sigma^2]''''') + 2[\sigma^2][\sigma^2]'''''' \right] + \mathcal{O}(P_L^3), \\ &= P_L + \int \frac{dk'}{2\pi} \left\{ 3 + 4\frac{k-k'}{k'} + \frac{(k-k')^2}{k'^2} \right\} P_L(k') P_L(k-k') + \dots, \end{aligned}$$

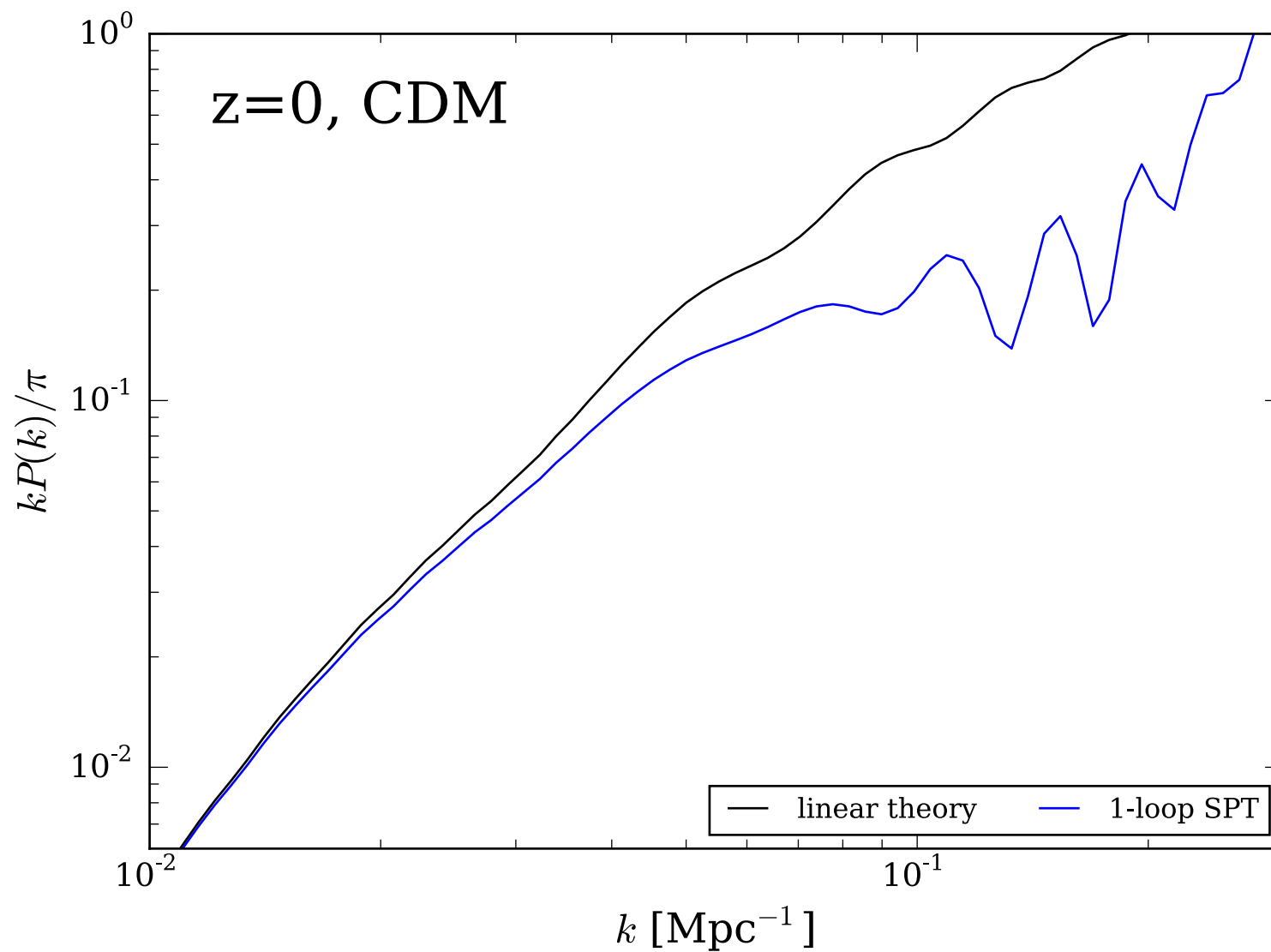
$$F_n(k_1, \dots, k_n) = G_n(k_1, \dots, k_n) = \frac{1}{n!} \frac{k^n}{\prod_{i=1}^n k_i}$$

Now we can understand common resummation schemes in “standard” perturbation theory, and we can look at the rate of convergence of perturbation theory.

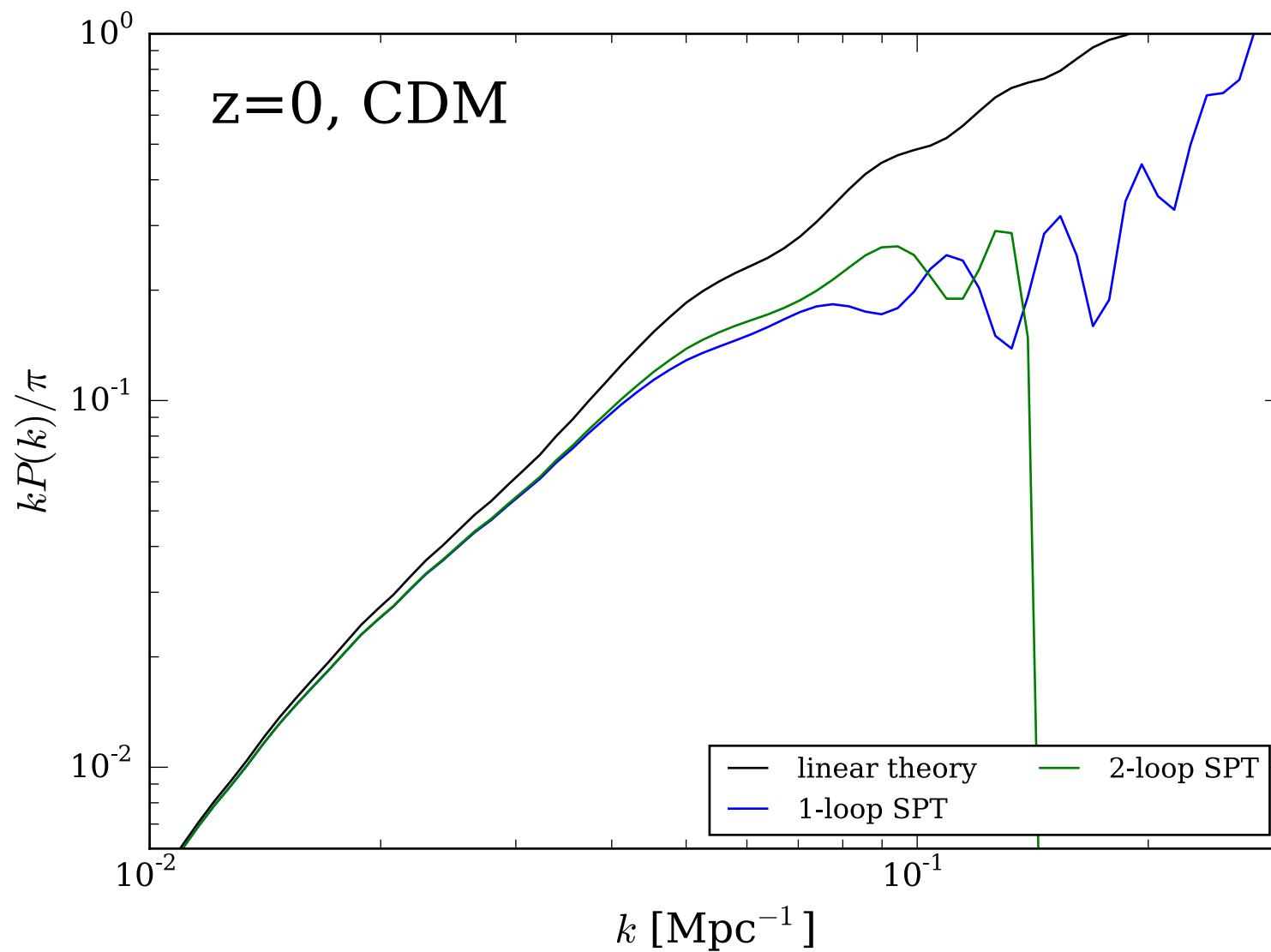
Convergence of PT



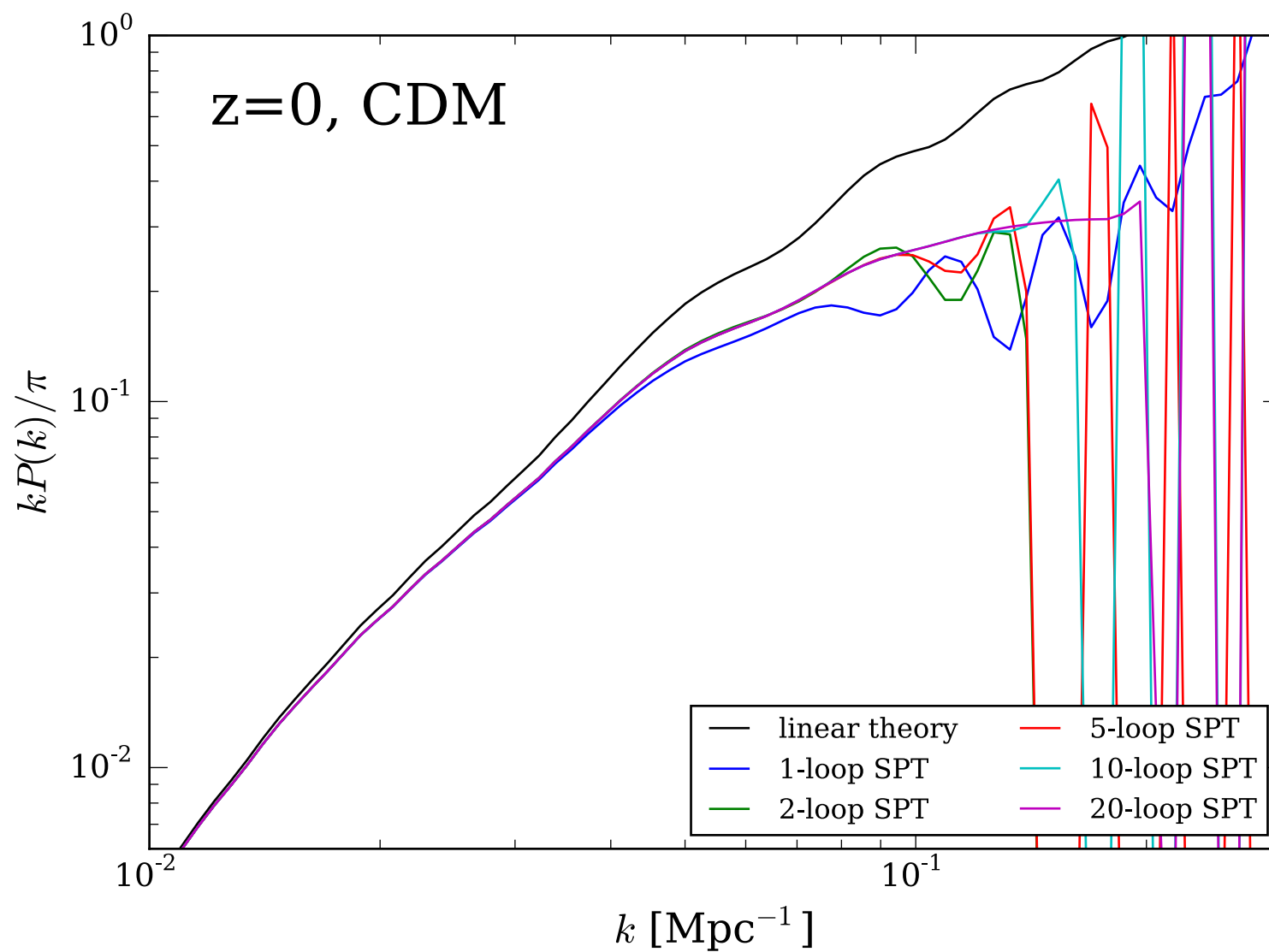
Convergence of PT



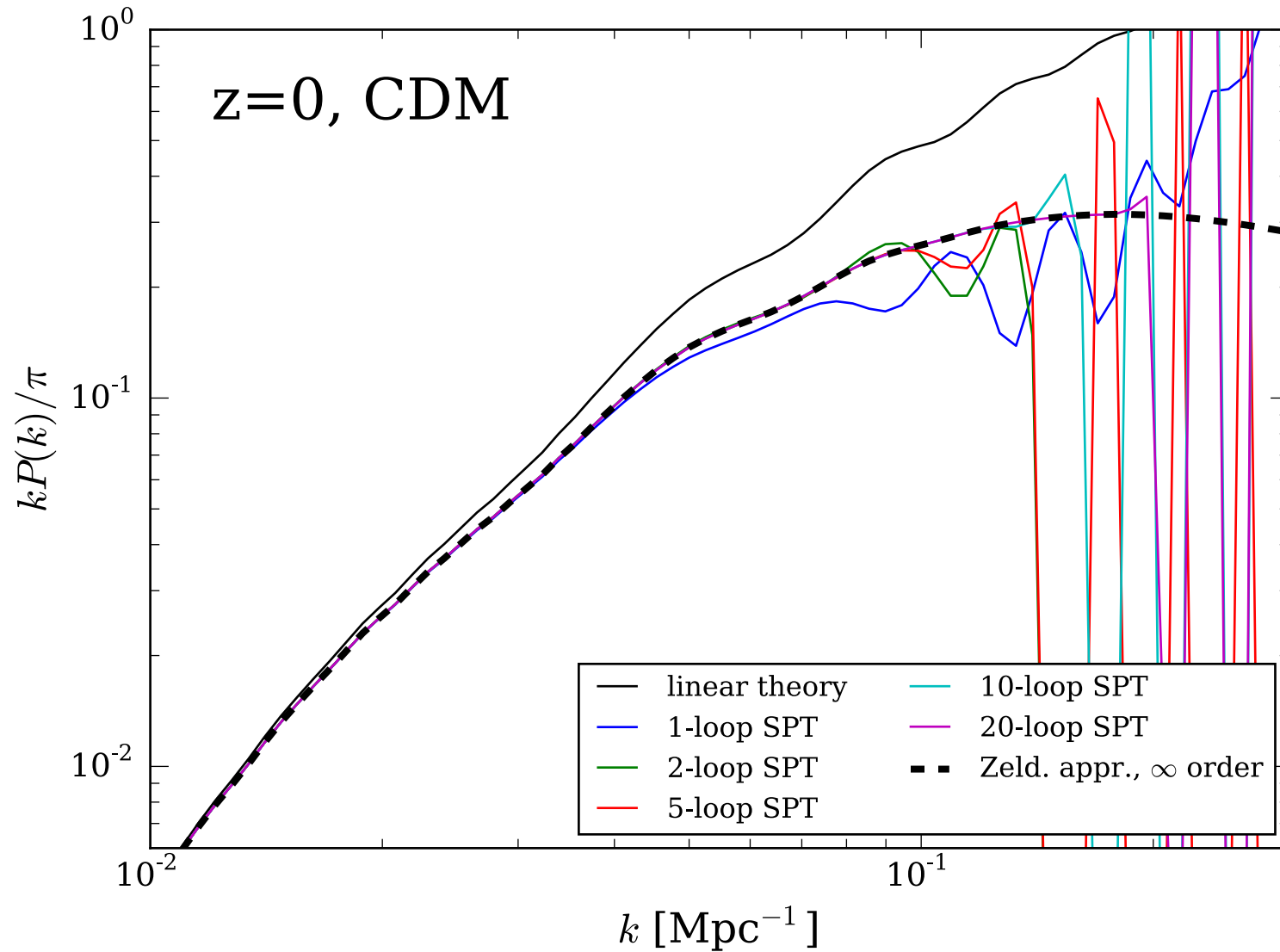
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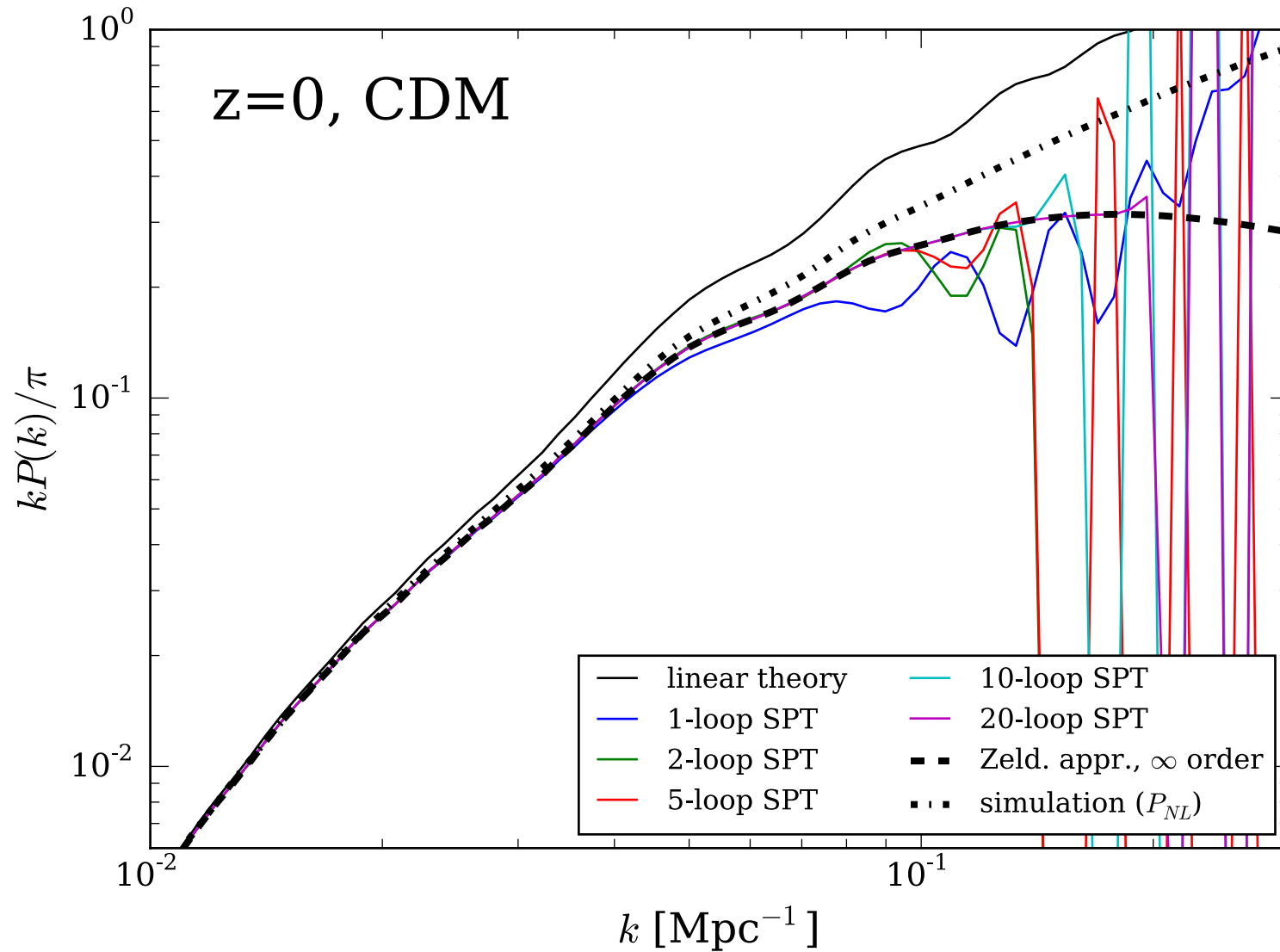
Convergence of PT

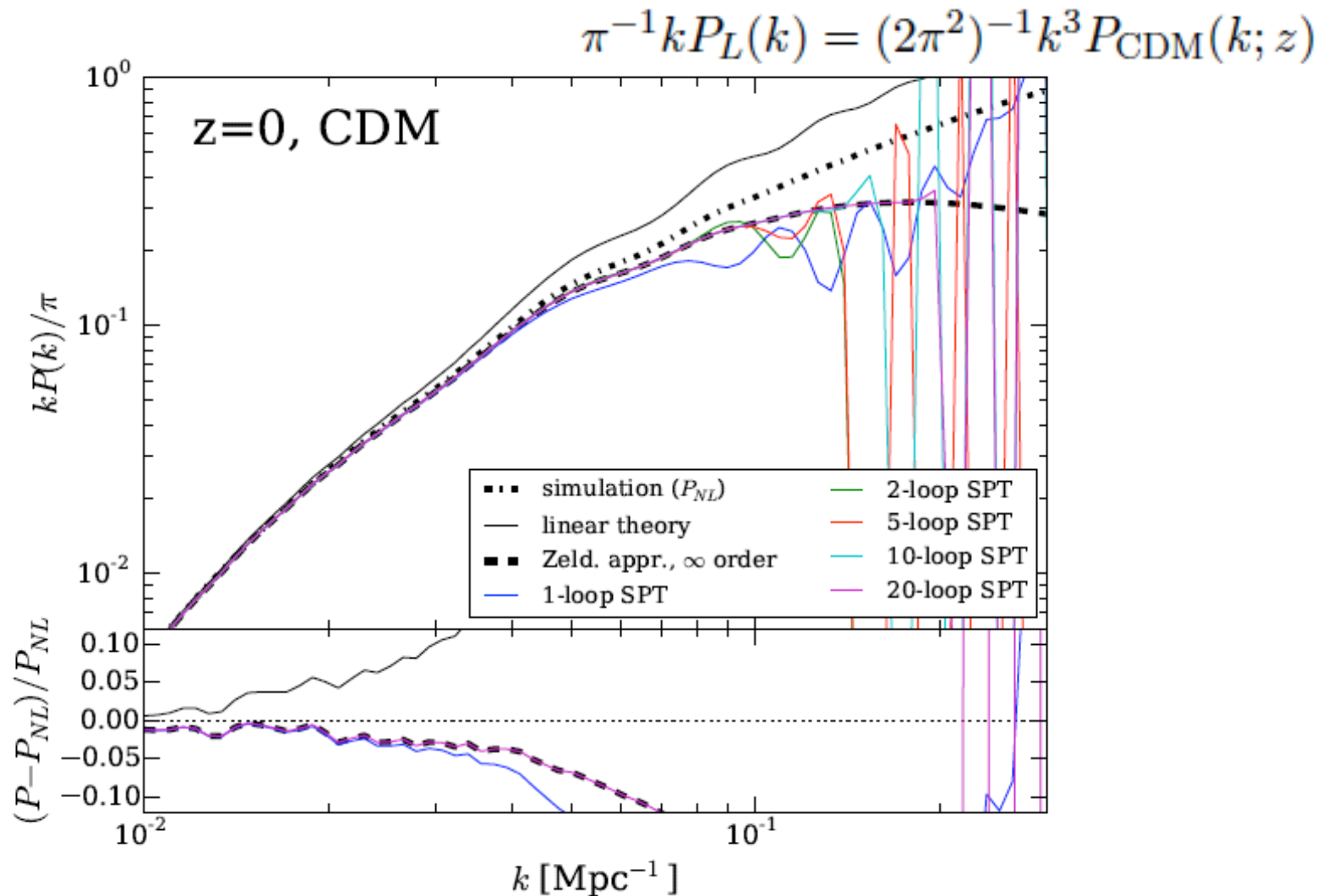


Convergence of PT



Convergence of PT





For 1D CDM-like cosmology, standard perturbation theories do not describe evolution on any non-linear scale accurately.

Some other results & future directions

- We have movies of the evolution of $\delta(x)$ which can be helpful for gaining intuition.
- Proof that SPT and LPT solutions identical to all orders, even though describe different physical systems!
- Easier to see the breaking of Galilean invariance that occurs in some schemes.
- Can formulate and test “effective field theory of large-scale structure” or “coarse grained perturbation theory”.
 - Correct and clarify some things in literature, give simpler derivations of some results.
 - A new Lagrangian EFT.
 - For power-law power spectra these theories get a lot of stuff “right”. Still some issues though ...
- Want to study bias/halos and redshift-space distortions.

The End